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# SHELL SPLITTING DUE TO ELECTRIC FIELDS

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Shell Splitting Due to Electric Fields

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Drift shells of trapped particles starting from a single field line in the earth's magnetic field may become separated due to an electric field. Using the formalism of the shell splitting function, this separation is calculated for an arbitrary curl-free time-independent electric field stisfying  $\underline{\mathbf{E}} \cdot \underline{\mathbf{B}} = \mathbf{0}$ , with no limitation on the particle's longitudinal excursion. The effect of non-dipole field components is also taken into account to the lowest order.

In a dipole field trapped particles starting from a common field line with different values of the mirror-point field intensity  $E_{g_1}$  will, in general, share the same drift shell. If the field suffers an asymmetric perturbation, this no longer occurs and the drift shells of particles with different  $E_{g_1}$  will separate as the distance from the starting field line increases. This effect is commonly known as shell splitting. It is of interest, since in conjunction with pitch-angle scattering it may transport particles across field lines.

Roederer and Schulz [1971], shell splitting due to a certain choice of electric field was derived, for particles with small longitudinal excursions (i.e. near-equatorial particles) in a dipole magnetic field.

Roederer and Schulz derived their result more or less from first principles, but a systematic approach also exists, based on the concept of the shell splitting function [Stern, 1968, henceforth referred to as A]. This approach will be used here to generalize the preceding result and to derive shell splitting due to a more general electric field. The calculation is not restricted to near-equatorial orbits and can be readily generalized for magnetic fields that deviate somewhat from the dipole configuration.

#### SHELLS IN A DIPOLE FIELD

Consider first only the dipole component  $\underline{B}_{0}$  of the earth's magnetic field; quantities pertaining to this component will henceforth be distinguished by subscript zero. This field may be described by Euler potentials  $(\prec_{0}, \ \ \ \ \ \ \ \ \ )$ 

$$\underline{B}_{0} = \nabla \alpha_{0} \times \nabla \beta_{0} \tag{1}$$

[Stern, 1970, and references listed there], with

$$\propto_{0}^{2} = a g_{1}^{0} (a/r) \sin^{2} \theta$$
 (2)

$$\beta_0 = a \Psi \tag{3}$$

Here a is the earth's radius,  $g_1^0$  is the axial dipole coefficient of the scalar magnetic potential and  $(r, \theta, \psi)$  are spherical dipole coordinates, so that  $g_1^0$  describes the entire dipole component of the field.

In the absence of electric fields the momentum p is conserved and the longitudinal invariant J becomes

$$J = 2 p I$$

$$= 2 p \int_{B_{m}^{1}}^{B_{m}} \left\{ 1 - B_{0}/B_{m} \right\}^{\frac{1}{2}} ds \qquad (4)$$

with the integration carried along a field line. To facilitate this integration in a dipole field one may express the field intensity as

$$I_{0}(\propto_{0}, B_{m}) = \int_{\pi-6m}^{\pi} \left\{1 - B_{0}(\propto_{0}, e) / B_{m}\right\}^{\frac{1}{2}} (\Im s/\sigma e) de$$
(6)

where s is differentiated with  $\sim_{O}$  held constant and where  $\cap_{m}$  is the colatitude associated with  $B_{m}$  (for the actual use of  $I_{O}$  and similar functions, see <u>Stern</u>, 1967a). The equation of the drift shell associated with a given pair of values of  $(I, B_{m})$  is then

$$I = I_0( \propto_0, B_m) \tag{7}$$

and this may be inverted to

$$\propto_{O} = G_{O}(I, B_{m}) \tag{8}$$

Note that this is a particular case of the general equation

$$f(\propto_0, \beta_0, I, B_m) = 0$$
 (9)

of a two-parameter family of surfaces that are everywhere orthogonal to  $\underline{B}_{O}$  in a dipole field using ( $\beta_{O}$  is absent, due to symmetry). Instead of, the parameters (I,  $B_{m}$ ) one can characterize this family by any two independent functions of these parameters. In particular, if  $G_{O}$  is one of these functions, the shell equation depends only on one parameter, namely  $G_{O}$ . Thus the function  $G_{O}$  is closely related to McIlwain's shell parameter L McIlwain, 1961 and in fact

$$L(I, B_m) = a g_1^O / G_O$$
 (10)

#### SHELL SPLITTING BY AN ELECTRIC FIELD

Let now a static electric field be added

$$\underline{\mathbf{E}} = - \nabla \phi \tag{11}$$

with no time dependence and no component parallel to  $\underline{\mathtt{B}}_{\mathrm{O}}$  . This last condition implies that

$$\phi = \phi ( \propto_0, \beta_0)$$
 (12)

Let it furthermore be assumed that this field is relatively weak, so that the motion of particles moving in the vicinity of the dipole is only slightly modified. For particles trapped in the combined electromagnetic field J and  $B_{\rm m}$  are still conserved, but I is not, since p is no longer a constant. One may nevertheless resolve J into a product involving I

$$J = p I \tag{13}$$

If at a particular instant the field line followed by the particle is characterized by Euler potentials ( $\propto_0$ ,  $\beta_0$ ), then I is expressed by the function I<sub>0</sub> of eq. (6) while p is given by

$$p^{2}c^{2} = W^{2} - m^{2}c^{4}$$

$$= (W_{s} + e[\phi(\propto_{0}, \beta_{0}) - \phi_{s}])^{2} - m^{2}c^{4}$$
(14)

Here W is the energy and the subscript s refers to some starting field

line to which all energies are referred (obvious choices would be field lines in the noon meridian or in the midnight meridian). Neglecting second order corrections (as will henceforth always be assumed)

$$p^{2}c^{2} = p_{s}^{2}c^{2} + 2eW_{s}[\phi(\alpha_{0}, \beta_{0}) - \phi_{s}]$$
 (15)

$$p = p_s \left\{ 1 + \left( e W_s / p_s^2 c^2 \right) \left[ \phi \left( \alpha_0, \beta_0 \right) - \phi_s \right] \right\}$$
 (16)

Let a constant  $I_s$  be defined, equal to I on the initial field line

$$I_{s} = J/p_{s} \tag{17}$$

Substituting (16) in (13) then gives as the shell equation to lowest order

$$I_{O}(\propto_{O}, B_{m}) = (1 - \delta) I_{s}$$
 (18)

where

$$\delta (p_s, \alpha_o, \beta_o) = (e W_s / p_s^2 c^2) \left[ \phi (\alpha_o, \beta_o) - \phi_s \right]$$
(19)

As in the transition between equations (7) and (8), this may be inverted to give the shell equation as

The second term on the right is what was defined in reference A as the "shell splitting function" and it distinguishes the shells in the present case from those given by eq. (8) for the case when the electric field is absent. To evaluate this function one notes that, since (8) is the inverse of (6)

$$I = I_0 \left[ G_0(I, B_m), B_m \right]$$
 (21)

Differentiation yields

$$\mathbf{1} = (\mathbf{r}_{1}/\mathbf{r}_{0})(\mathbf{r}_{0}/\mathbf{r}_{1}) \tag{22}$$

and hence

$$(\Im G_0/\Im I) = (\Im I_0/\Im \omega_0)^{-1}$$
 (23)

It may be shown [Stern, 1967a] that

where  $K_4$  is an integral defined by <u>Pennington</u> [1961]. The expression given in (24) does depend on  $\bowtie_0$ , but since  $\bowtie_0$  varies only slightly around a drift shell, we are justified (at least to lowest order) in replacing it with its initial value  $\bowtie_s$ , and a similar substitution may be performed in  $\diamondsuit$ . This gives the shell equation to lowest order as

$$\propto 0 = G_0(I_s, B_m) + (2 \propto_s^2 \cdot W_s / a^2 g_1^0 K_4 c^2 p^2) [\phi(\propto_s, \beta_0) - \phi_s]$$

The second term in (25) gives explicitely the effect of the electric field on the shape of drift shells. Since it depends on both  $W_s$  and  $\beta_0$ , it indeed leads to energy-dependent shell splitting.

#### SHELL SPLITTING IN NEAR - DIPOLE FIELDS

In slightly asymmetrical fields the equations of drift shells are further modified in the manner described in reference A. To the first order the two modifications may be independently applied, leading to a first-order shell equation

where  $\propto_1$  is the first-order correction to the Euler potential  $\begin{bmatrix} \text{Sterm, 1967b} \end{bmatrix}$  and where  $G_1$  is the shell splitting function for near-dipole magnetic fields, as given in A and also as implicit in the work of Pennington [1961]. The last two terms both lead to shell splitting, since both contain  $\frac{3}{100}$ , but only the latter one includes dependence on the initial energy  $W_8$ .

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#### REFERENCES

Chapman, S., Geomagnetic Nomenclature, J. Geophys. Res. 68, 1174, 1963.

- McIlwain, C.E., Coordinates for Mapping the Distribution of Magnetically Trapped Particles, J. Geophys. Res. 66, 3681, 1961.
- Pennington, R.H., Equation of a Charged Particle Shell in a Perturbed Dipole Field, <u>J. Geophys. Res.</u> 66, 709, 1961.
- Roederer, J.G. and Schulz, M., Splitting of Drift Shells by the Magneto-spheric Electric Field, <u>J. Geophys. Res.</u> 76, 1055, 1971.
- Stern, D.P., Euler Potentials and Geomagnetic Drift Shells, Goddard Space

  Flight Center Document X-641-67-329, 1967a.
- Stern, D.P., Geomagnetic Euler Potentials, J. Geophys. Res. 72, 3995, 1967b .
- Stern, D.P., Euler Potentials and Geomagnetic Drift Shells, <u>J. Geophys. Res.</u>

  73, 4373, 1968.
- Stern, D.P., Euler Potentials, American J. of Phys. 38, 494, 1970 .